Human Initiated Contact — Jaya Sanctuary

MUPAS Sensor Data Analysis Report — with Baseline Control Window

Jaya Sanctuary, Victoria, Australia | September 12, 2026 | Event window 1:00 PM – 8:30 PM local

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1. Executive Summary

This report presents a scientific analysis of multi-channel environmental sensor data recorded during a Human Initiated Contact session held at Jaya Sanctuary in Victoria, Australia on September 12, 2026. The event ran from 1:00 PM to 8:30 PM local time under clear skies with light cloud cover, a new moon, and temperatures falling from 21°C to around 10°C. Fourteen witnesses were present. The session was hosted and led by Peter Maxwell Slattery, assisted by Jaya team member Julie, using the established protocol of Heart Breathing for group coherence, Clearing the Space, and Initiating Contact meditations repeated throughout the event, with the principal Initiating Contact meditation completed at 6:20 PM. The first sighting — an orb, captured in a snapshot — occurred at 6:24 PM, immediately after that meditation and coinciding with disturbances on the MUPAS portal. One UAP "Power Up" response was logged at 7:21 PM. As at prior events, everything logged as anomalous was absent from the satellite tracker, while normal aircraft flight paths were separately seen and acknowledged. Geomagnetic conditions were quiet (G0, no storms, B-class solar flares only), giving a clean interpretation context for the magnetic channel.

Like the October 17 event, this session was recorded with a same-day baseline control window (8:30 AM – 12:30 PM) in addition to the session window (charted here 1:30 – 5:30 PM). The headline finding is a replication: the session pressure trace again contains patches of rapid, noisy flickering — concentrated near 2:55 – 3:05 PM and 3:40 – 3:50 PM — while the same instrument's baseline pressure trace from the same morning is perfectly smooth for four straight hours. This is now the second independent baseline-controlled event at this site in which the anomalous pressure noise appears only during the session window. The session record also contains two multi-channel co-occurrence clusters: pressure noise + infrasound burst + paired large microwave spikes at ~3:40 – 3:50 PM, and a magnetometer excursion + gyroscope spike + accelerometer dip cluster at ~5:10 – 5:25 PM.

2. Event Overview and Observation Log

2.1 Site, Conditions, and Equipment

Jaya Sanctuary, in the Carlyle/Rutherglen area of Victoria (36.3°S, 145.9°E), hosted the session under clear to mostly clear skies with light southwest winds and a new moon. Equipment deployed included a SIONYX Aurora PRO night vision camera, a Sony A7s2 camera with 28–70mm lens, the MUPAS unit, and the Night Sky satellite tracker for real-time conventional-object exclusion.

2.2 Space Weather Context

The space weather summary for September 12–13 records geomagnetic activity at G0 (quiet), with no storms or major disturbances, very low solar flare activity (mostly B-class, no significant threats), and no aurora watch or alert for the location. As with the October 17 event — and unlike the May 16 event, which took place during a Kp 7–8 storm — any structure in the magnetometer record on this night cannot be attributed to geomagnetic storm activity.

2.3 Observation Timeline

1:00 PM — Session begins; MUPAS recording.

6:20 PM — Initiating Contact meditation completed.

6:24 PM — First sighting: an orb, photographed, appearing immediately after the IC meditation, coinciding with readings on the MUPAS portal (Section 6).

7:21 PM — One UAP "Power Up" response logged. The portal record shows the session's largest RF event (~27,300 units at 7:12:16 PM) nine minutes before this response, with a magnetometer baseline step between the two (Section 6).

8:30 PM — Session ends.

3. The Baseline Control Window

A four-hour baseline (8:30 AM – 12:30 PM local) was recorded with the same MUPAS unit at the same location on the morning of the event, before participants assembled. As established in the October 17 analysis, the control logic is: a signature present in both windows is site or instrument background; a signature present only in the session window — particularly one aligning with documented activity — is a genuine session-specific anomaly. September 12 chronologically precedes October 17, so the programme now holds two independent baseline-controlled events at this site, recorded five weeks apart, which allows replication testing — the strongest standard of evidence this programme has yet been able to apply.

4. Sensor Data Record

Figure 1 presents the morning baseline; Figure 2 presents the charted session window (1:30 – 5:30 PM). The evening observation period (6:20 PM meditation, 6:24 PM orb, 7:21 PM Power Up) falls after the charted window and is covered narratively from the portal captures in Section 6.

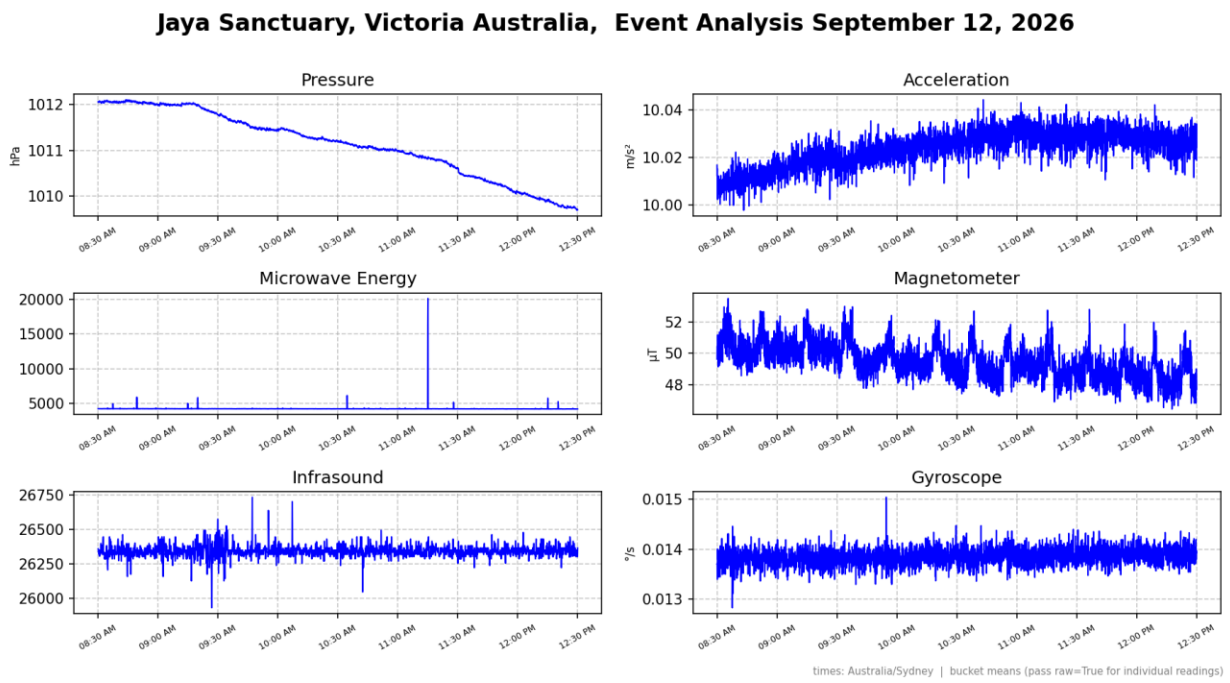


Figure 1: BASELINE — Six-channel MUPAS record, Jaya Sanctuary, September 12, 2026, 8:30 AM – 12:30 PM local (Australia/Sydney), recorded the same morning with the same instrument at the same location. Top row: Pressure (left), Acceleration (right). Middle row: Microwave Energy (left), Magnetometer (right). Bottom row: Infrasound (left), Gyroscope (right).

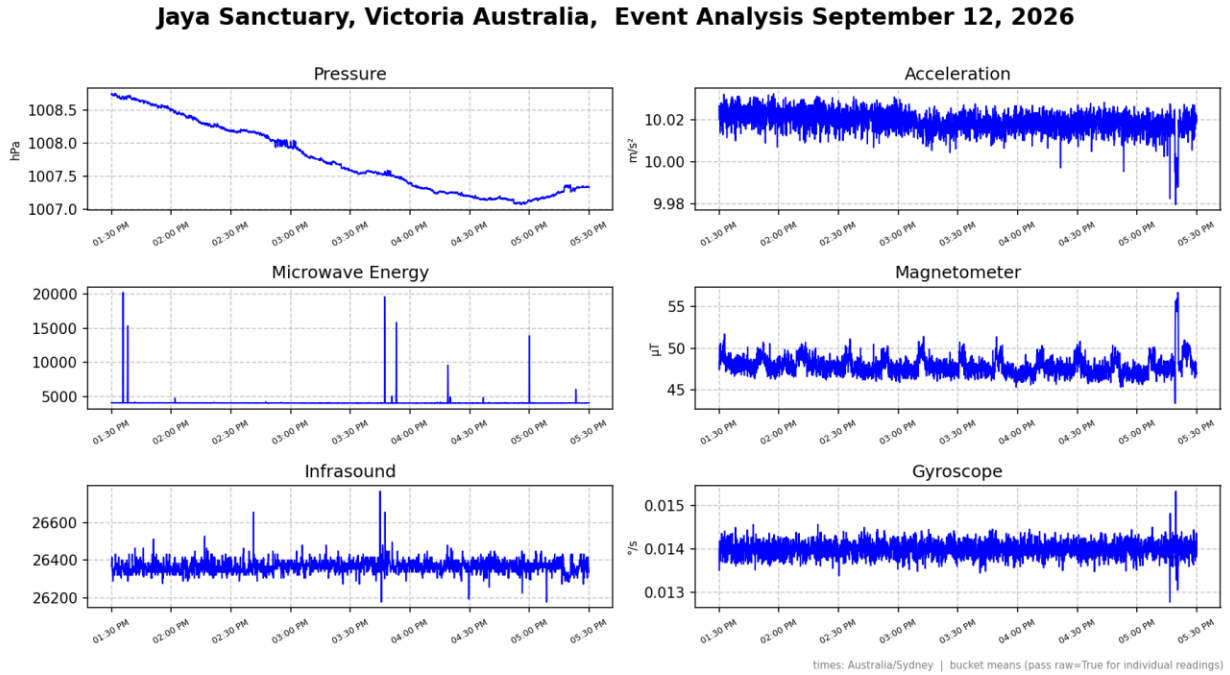


Figure 2: SESSION — Six-channel MUPAS record, same day and location, 1:30 PM – 5:30 PM local, covering the afternoon hours of the Human Initiated Contact session. Same panel layout as Figure 1.

5. Anomaly Analysis — Session Against Baseline

5.1 Pressure — The Session-Only Noise Replicates

The baseline pressure trace is immaculate: a smooth decline from ~1012.05 to ~1009.7 hPa over four hours (~2.3 hPa of ordinary synoptic fall ahead of an approaching trough), without a single discontinuity, step, or noisy interval — four straight hours of textbook barometric behaviour from the very instrument in question, on the very day in question.

The session trace declines over a similar slope (~1008.7 to a floor near 1007.1 hPa by ~5:00 PM, then an upturn) — but is interrupted by patches in which the trace abruptly stops behaving like a barometric record and becomes visibly noisy: a fuzz of rapid, small, repeated drop-and-recovery flickering superimposed on the smooth curve. The two clearest patches sit at approximately 2:55 – 3:05 PM (around the 1008.0 hPa level) and 3:40 – 3:50 PM (around 1007.6 hPa), each lasting on the order of ten minutes before the trace reverts to smooth behaviour, with a further brief disturbed interval near 3:55 PM.

This is the same signature isolated at the October 17 event, where the mid-session notch cluster and ~5:30 PM flickering appeared against a smooth same-day baseline. The replication is the point: two independent events, five weeks apart, same site, same instrument, each with its own same-day control — and in both, the rapid pressure flickering occurs only during the session hours. An instrument defect would not respect the session schedule twice; a weather process would have to afflict only afternoons on which a contact session happened to be running. Atmospheric pressure reflects the weight of the entire air column overhead and cannot naturally flicker on a timescale of seconds, repeatedly, and then revert — so each patch implies either something transiently and repeatedly disturbing the air in the sensor's immediate vicinity, or an external influence acting on the sensing element itself. With two controlled replications, the session-only pressure noise is now the best-evidenced instrumental signature in this research programme.

5.2 The ~3:40 – 3:50 PM Multi-Channel Cluster

The second pressure-noise patch does not occur alone. Within the same ten-minute window the record shows: two large microwave spikes (~19,500 and ~16,000 units, the session's second- and third-largest, at ~3:40 and ~3:50 PM) and the session's most active infrasound interval — a burst centred near 3:40 – 3:45 PM reaching ~26,780 units upward and ~26,180 downward, the largest excursions in either direction in the charted window.

Per the October 17 baseline finding, microwave spike amplitude alone is not probative at this site — the September 12 baseline likewise contains a ~20,000-unit spike on a quiet morning, confirming that large sporadic RF bursts are site background. What the baseline logic cannot explain away is timing: the probability that independent sporadic background events on three channels happen to land inside the same ten-minute window that also contains the session-only pressure anomaly is low, and this triple co-occurrence pattern — pressure, infrasound, and electromagnetic channels moving together — is the programme's established central signature, here appearing in a baseline-controlled record for the first time with the pressure channel as a participant.

5.3 The ~5:10 – 5:25 PM IMU/Magnetometer Cluster

Near the end of the charted window a second, different cluster appears. The magnetometer — which to that point had varied within a ~45.5 – 51 µT envelope similar to baseline — produces its largest event of the window at ~5:15 – 5:20 PM: a sharp excursion rising to ~56.5 µT and plunging to ~43.5 µT, a ~13 µT peak-to-trough swing far exceeding anything in the baseline. Simultaneously, the gyroscope registers its largest spike of the window (~0.0153 deg/s against

a 0.0140 floor), and the accelerometer — whose isolated one-sided dips are known site background — departs from that background pattern by producing a dense cluster of deep dips (to ~9.98 m/s²) concentrated in the same ~5:05 – 5:25 PM interval rather than scattered randomly.

Under G0 quiet geomagnetic conditions, a ~13 µT local magnetic swing has no space weather explanation; it requires a strong local field source or disturbance at the sensor. Its synchrony with the gyroscope's only notable event and the accelerometer's only clustered behaviour marks this as a genuine multi-channel IMU/magnetic event — the same family as the Ozarks MagZ/GyroZ anomaly and the October 17 ~8:00 PM portal step, and now the third appearance of this signature class in the series.

5.4 Channels Consistent with Background

For completeness and discipline: the session's early microwave spikes (~20,000 and ~15,500 units at ~1:35 – 1:40 PM) and remaining isolated RF events (~9,700 at ~4:15 PM; ~14,000 at ~5:00 PM) are treated as site-characteristic background per the baseline standard, noted only where their timing coincides with other channels. The baseline's own infrasound record shows an active interval (~9:20 – 10:10 AM) of moderate spikes, so infrasound incidence alone is likewise not probative at this site — the session's 3:40 PM burst is claimed only as part of the timed cluster in Section 5.2. The magnetometer's general variability envelope and the scattered single accelerometer dips match baseline behaviour and are claimed as nothing. This is the baseline methodology working as intended: the analysis rests only on what the control cannot explain.

6. The Evening Window — Meditation, Orb, and Power Up

The documented visual events fall after the charted session window and are evidenced by the portal captures included in the field report (device mupasm-0524-15). Three portal observations are salient:

> ~6:20 – 6:30 PM — Immediately following completion of the Initiating Contact meditation at 6:20 PM, and coinciding with the 6:24 PM orb photograph, the portal's 3:30 – 6:30 PM capture shows a simultaneous disturbance across the magnetometer and gyroscope axes at the right edge of the window (~6:20 – 6:30 PM): an abrupt coordinated transient in MagX/MagY/MagZ with concurrent gyroscope disturbance — the same coordinated Mag/Gyro signature class as the 5:10 – 5:25 PM cluster an hour earlier, and the same post-meditation timing pattern noted at both prior Jaya events.
>
> 7:12:16 PM — The portal records the evening's largest RF event, ~27,300 units (portal tooltip), roughly nine minutes before the logged 7:21 PM Power Up response.
>
> ~7:15 PM — Between the RF event and the Power Up, the magnetometer baseline undergoes a clean step change (visible as a level shift in the MagX/MagZ traces) that persists for the remainder of the evening record.

The sequence — contact meditation completed, orb photographed four minutes later with coordinated Mag/Gyro disturbance, then the evening's largest RF burst, a persistent magnetometer level shift, and a witnessed Power Up, all within the following hour — is the most tightly time-ordered protocol-to-phenomenon-to-sensor chain yet documented at this site. Extracting the full-resolution 6:00 – 8:30 PM record and analysing it at chart level against the same-day baseline is the top follow-up for this case.

7. Comparison with Prior Events in This Research Programme

Six prior field events have been analysed in this series: the Ozarks CE6 (April 11–12, 2026), the CE6 Event of May 23, 2026, the Colorado CE6 (August 1, 2026), the Utah CE5 (September 10, 2026), and the Jaya Sanctuary Human Initiated Contact sessions of May 16 and October 17, 2026. September 12 is the third Jaya event analysed (chronologically the second held), making Jaya the programme's first site with three instrumented events — and, with October 17, its first replicated baseline-controlled pair.

7.1 Cross-Event Similarities

The pressure channel departs from quiet behaviour during the active window for the seventh event in seven: the Ozarks impulsive jump, the May 23 oscillations, the Colorado bowl, the Utah accelerating decline, the May 16 decline with slope change, the October 17 notch cluster and 5:30 PM flickering, and now the September 12 noise patches — the last two demonstrated against same-day controls that show nothing.

Post-meditation onset: the first sighting (6:24 PM) came immediately after the Initiating Contact meditation, with coordinated sensor disturbance at the same moment — the third consecutive Jaya event with this pattern, echoing protocol-milestone alignment at the Ozarks and Colorado.

Multi-channel co-occurrence: two distinct clusters in the charted window (pressure/infrasound/RF at ~3:40 – 3:50 PM; Mag/Gyro/Accel at ~5:10 – 5:25 PM) plus the evening Mag/Gyro transient — continuing the series' central diagnostic across a baseline-controlled record.

Real-time conventional exclusion: anomalous activity absent from the satellite tracker while normal flight paths were seen and acknowledged — consistent across every Jaya and Utah event.

7.2 What Is Distinctive About September 12

First replication of the session-only pressure noise: October 17 isolated the signature; September 12 shows the same instrument producing the same session-only flickering against its own smooth same-day baseline, five weeks apart. A one-off anomaly has become a repeatable, controlled observation — a categorical upgrade in evidential standing.

Largest controlled magnetometer event: the ~13 μT swing at ~5:15 PM under G0 quiet conditions exceeds any magnetic excursion in the series outside the Ozarks, and is the first with both a quiet-sky context and a same-day control.

The timed evening chain: meditation → orb (4 minutes) → largest RF burst → persistent magnetometer step → witnessed Power Up (9 minutes after the RF burst) is the series' most granular time-ordered sequence linking protocol, observation, and instrumentation.

Background discipline extended: the baseline reconfirmed large microwave spikes as site background and newly flagged that infrasound incidence alone is non-probative at this site — further narrowing, and further hardening, the programme's evidential core.

7.3 Updated Cross-Event Picture

With seven analysed events, the programme's surviving core is: (1) a pressure channel that misbehaves during every active session and never during controls — now replicated twice

under same-day baseline conditions at one site; (2) multi-channel co-occurrence clusters at documented moments, appearing in every event including baseline-controlled ones; and (3) protocol-milestone timing alignments, now including a nine-minute RF-to-Power-Up sequence. Amplitude-based claims (microwave, isolated accelerometer dips, infrasound incidence) have been progressively retired by the baseline methodology. What remains is smaller, harder, and — for the first time — replicated.

8. Summary and Recommendations

The September 12, 2026 Human Initiated Contact session at Jaya Sanctuary produced: session-only pressure noise patches (~2:55 – 3:05 and ~3:40 – 3:50 PM) replicating the October 17 signature against a flawless same-day baseline; a triple co-occurrence cluster (pressure noise, infrasound burst, paired large RF spikes) at ~3:40 – 3:50 PM; a ~13 µT magnetometer excursion with synchronous gyroscope spike and accelerometer dip cluster at ~5:10 – 5:25 PM under quiet geomagnetic conditions; and a time-ordered evening chain from the 6:20 PM contact meditation through the 6:24 PM photographed orb, coordinated Mag/Gyro disturbance, the 7:12 PM ~27,300-unit RF burst, a persistent magnetometer step, and the 7:21 PM witnessed Power Up.

Recommended follow-up actions:

- Extract and analyse the full-resolution 6:00 – 8:30 PM MUPAS record at chart level, covering the orb, the RF burst, the magnetometer step, and the Power Up, against the same-day baseline.
- Characterise the two pressure-noise patches at raw sample resolution (amplitude distribution, inter-flicker timing) and compare quantitatively against the October 17 notch cluster to test whether the two events share one signature.
- Investigate the persistent ~7:15 PM magnetometer step — determine whether the device was moved or whether the level shift is environmental.
- Continue the same-day baseline as standard practice, and add an evening-hours baseline on a non-event day to control for diurnal effects at this site.
- Maintain the practice of logging protocol milestones with clock times; the 6:20 PM meditation timestamp is what makes the evening chain analysable.

This report is a scientific analysis of sensor data, event records, and space weather context as supplied. Conclusions are interpretive and intended to guide further investigation, not to constitute definitive claims about the nature of the observed phenomena.